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## Gaseous HCN<sup>+</sup>, HNC<sup>+</sup>, and HCNH<sup>+</sup> Ions<sup>1</sup>

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**Abstract:** Two isomers of CHN<sup>+</sup> can be identified from their collisional activation mass spectra. Molecular HCN ions are predominantly the HCN<sup>+</sup> isomer, but decomposition of a wide variety of N-organic ions produces only the more stable HNC<sup>+</sup> isomer. Of the three possible isomers of CH<sub>2</sub>N<sup>+</sup>, only HCNH<sup>+</sup> is found to be produced by a variety of decomposition and ion-molecule reactions. Formation of these ions in the mass spectra of CH<sub>3</sub>CN and CH<sub>3</sub>NC is unusual in involving the neutral loss of CH<sub>2</sub> and CH, which apparently is due to the large stabilization resulting from the addition of one and two hydrogens to CN<sup>+</sup>.

Molecular HCN is stable, while its isomer HNC has been observed only as a transient species.<sup>2</sup> In contrast, the HNC<sup>+</sup> ion is more stable than its HCN<sup>+</sup> isomer by 1.3 eV according to theoretical calculations,<sup>3</sup> and by 0.6 eV from charge exchange experiments.<sup>4</sup> In the latter ground-state formation of the isomers HCN<sup>+</sup> from HCN ionization and HNC<sup>+</sup> by unimolecular decomposition of CH<sub>3</sub>NC<sup>+</sup> and CH<sub>3</sub>NH<sub>2</sub><sup>+</sup> was assumed. We report here a further investigation utilizing collisional activation (CA) mass spectra<sup>5</sup> of the formation of these CHN<sup>+</sup> isomers over a range of energies from a variety of precursors. Because the isobaric C<sub>2</sub>H<sub>3</sub><sup>+</sup> (*m/z* 27) ions commonly occur in mass spectra from background as well as fragmentation, this study utilized a CA instrument capable of high resolution separation of such interfering ions.<sup>6</sup>

The reported<sup>4</sup> formation of HNC<sup>+</sup> from CH<sub>3</sub>NC is of particular interest because this must involve the loss of CH<sub>2</sub>, a neutral loss that is highly unusual in mass spectra.<sup>7</sup> Our initial examination of the mass spectrum of CH<sub>3</sub>NC not only confirmed the formation of CHN<sup>+</sup> (2% of base peak), but also showed a more abundant CH<sub>2</sub>N<sup>+</sup> peak (3%) formed by the even more surprising loss of CH. Further, we find that CH<sub>3</sub>CN<sup>+</sup> also shows similar losses of CH as well as CH<sub>2</sub>. CH<sub>2</sub>N<sup>+</sup> ions have attracted substantial interest with reference to their role in the interstellar formation of HCN and HNC. Although H<sub>2</sub>CN<sup>+</sup> has been suggested as the stable

isomeric structure,<sup>8</sup> theoretical calculations by Schaefer<sup>9</sup> predict that the linear HCNH<sup>+</sup> is the most stable isomer, with CNH<sub>2</sub><sup>+</sup> and H<sub>2</sub>CN<sup>+</sup> being 2.0 and 3.1 eV, respectively, less stable.

### Experimental Section

The tandem CA mass spectrometer<sup>6</sup> used has a high-resolution Hitachi RMH-2 as MS-I. Ions formed by 70-eV electrons or by chemical ionization (CI) are accelerated by 10 keV, separated by MS-I, and made to undergo collisions with a helium molecular beam; the resulting CA product ions are separated by an electrostatic analyzer used as MS-II. The helium flow rate was increased until the precursor ion abundance was reduced to 40% of its original value; at an equivalent pressure, the abundance of a precursor ion such as C<sub>2</sub>H<sub>3</sub><sup>+</sup> would be reduced to ~10% of its original value. Spectra were computer averages of ~20 scans; the data for CH<sub>2</sub>N<sup>+</sup> from CH<sub>3</sub>NC and CH<sub>3</sub>CN are averages of ≥5 such runs. Confirmatory data were also measured on a reverse-geometry RMU-7 double-focusing mass spectrometer<sup>10</sup> using 8-keV precursor ions, but only representative CA data for CHN<sup>+</sup>, corrected for C<sub>2</sub>H<sub>3</sub><sup>+</sup> interference, are reported.

The RMH-2-separated precursor CHN<sup>+</sup> ions (*m/z* 27.0109) contained less than 2% C<sub>2</sub>H<sub>3</sub><sup>+</sup> ions (27.0235) based on *m/z* 25 in the CA spectrum. Measuring the CA spectra of CH<sub>2</sub>N<sup>+</sup> (*m/z* 28.0187) required MS-I separation of H<sup>13</sup>CN<sup>+</sup> (28.0143), which differs by 1/6400, as well as of N<sub>2</sub><sup>+</sup> (28.0061) and CO<sup>+</sup> (27.9949). Care had to be taken to avoid drift of MS-I off the CH<sub>2</sub>N<sup>+</sup> peak during the multiple scanning of its CA spectrum; H<sup>13</sup>CN<sup>+</sup> contamination increases [*m/z* 14<sup>+</sup>]/[12<sup>+</sup>]. The CHDN<sup>+</sup> precursor ion was separated from CD<sup>15</sup>N, <sup>13</sup>CDN<sup>+</sup>, C<sup>13</sup>CD<sub>2</sub><sup>+</sup>, and C<sub>2</sub>HD<sub>2</sub><sup>+</sup>.

Protonation used to prepare precursor ions for reported CA spectra was done with H<sub>2</sub>O as the CI reagent gas. Self-protonation of HCN gave the same CA spectrum of CH<sub>2</sub>N<sup>+</sup> within experimental error.

Appearance energies, measured by the vanishing current method, were of poor accuracy because of the low ion abundance at high resolution. The CN<sup>+</sup> data also tallied badly.

Methyl isocyanide was synthesized by a standard literature procedure,<sup>11</sup> and the other compounds were obtained commercially.

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Table I. Collisional Activation Spectra of C(H,D)<sup>+</sup>N<sup>+</sup> Ions

compd	m/z					
	12	13	14	15	16	26
HCN	2.7	2.7	1.5	0.4		100
<i>a</i>	2.7	3.0	1.4	0.5		<i>b</i>
20 eV <sup>a,c</sup>	2.7	3.2	1.4	0.5		<i>b</i>
DCN	4.0		7.2		0.9	100
CH <sub>3</sub> NC	3.7	1.9	1.9	1.5		100
<i>a</i>	3.7	1.5	1.5	1.3		<i>b</i>
CH <sub>3</sub> NH <sub>2</sub> <sup>a</sup>	3.7	1.2	1.3	1.4		<i>b</i>
HCONH <sub>2</sub> <sup>a</sup>	3.7	1.3	1.4	1.6		<i>b</i>
CH <sub>3</sub> CN	3.6	2.0	1.5	1.2		100
HCON(CH <sub>3</sub> ) <sub>2</sub>	3.7	1.7	1.5	1.3		100
pyridine	3.9	2.3	1.6	1.1		100
benzonitrile	3.7	2.2	1.9	0.9		100

<sup>a</sup> Data from RMU-7; other data on the tandem/RMH-2 instrument. <sup>b</sup> RMU-7 determination of [12<sup>+</sup>]/[26<sup>+</sup>] inaccurate, so [12<sup>+</sup>] value of tandem instrument assumed. <sup>c</sup> Ionizing electron energy 70 eV unless otherwise specified.

Table II. Collisional Activation Spectra of C(H,D)<sub>2</sub>N<sup>+</sup> Ions

compd	m/z									
	12	13	14	15	16	17	18	26	27	28
HCN + H <sup>+</sup>	2.1	1.5	1.4	1.2	0.4			22	100	
HCN + D <sup>+</sup> <sup>a</sup>	1.7	1.0	1.2	0.2	1.5	0.4		19	36	67
CH <sub>3</sub> CN + D <sup>+</sup> <sup>a</sup>	1.7	1.2	1.4	0.6	1.0	0.2		13	32	77
DCN + H <sup>+</sup> <sup>a</sup>	1.3	0.4	2.3	0.8	0.2	0.2		17	42	62
CD <sub>3</sub> CN + H <sup>+</sup> <sup>a</sup>	1.6	0.5	2.3	1.1	0.8	0.3		20	40	62
DCN + D <sup>+</sup>	1.8		3.8	0.8 <sup>b</sup>	2.2		0.4	18		100
CH <sub>3</sub> NC	2.1	1.6	1.6	1.3	0.4			24	100	
CH <sub>3</sub> CN	1.9	1.5	1.7	1.3	0.3			21	100	

<sup>a</sup> Using CHDN<sup>+</sup> as the precursor ion. Abundances are normalized to the base [CN<sup>+</sup>] + [C(H,D)<sub>2</sub>N<sup>+</sup>] = 122; [CHDN<sup>2+</sup>] = 0.8. <sup>b</sup> CD<sub>2</sub>N<sup>2+</sup>.

## Results and Discussion

In confirmation of the results obtained by charge exchange,<sup>4</sup> the CA spectra of CHN<sup>+</sup> ions from methyl isocyanide and methylamine are the same, and differ significantly from the spectrum of CHN<sup>+</sup> ions from HCN (Table I). These spectra are consistent with the structures HNC<sup>+</sup> and HCN<sup>+</sup>, respectively, as the spectrum of the former shows more abundant HN<sup>+</sup> ions, and that of the latter has more abundant HC<sup>+</sup> ions. Assuming equivalent CA cross sections, [HN<sup>+</sup>] indicates that the molecular ions from HCN contain <30% of the HNC<sup>+</sup> isomer; however, the true value is probably much less than this, as HN<sup>+</sup> is probably formed from HCN<sup>+</sup> by rearrangement in the same way that HC<sup>+</sup> is formed from HNC<sup>+</sup>. Also lowering the ionizing electron energy does not change the CA spectrum from HCN appreciably.

The HNC<sup>+</sup> isomer is clearly the more stable.<sup>3,4</sup> This is the only isomer observed (Table I) for CHN<sup>+</sup> ions formed by mass spectral fragmentation of CH<sub>3</sub>NH<sub>2</sub>, CH<sub>3</sub>NC, CH<sub>3</sub>CN, HCONH<sub>2</sub>, HCON(CH<sub>3</sub>)<sub>2</sub>, pyridine, and benzonitrile. However, the energy barrier for the exothermic isomerization HCN<sup>+</sup> → HNC<sup>+</sup> can only be determined qualitatively from these data. Although for HCN ionization the appearance energy of the lowest energy dissociation process (H loss) is 5.8 eV above the ionization energy,<sup>12</sup> the photoelectron spectrum<sup>13</sup> indicates that >90% of the molecular ions are formed with <0.5 eV internal energy, so that this is the highest minimum value which can be derived for the

HCN<sup>+</sup> → HNC<sup>+</sup> barrier from these data.

**CH<sub>2</sub>N<sup>+</sup> Ions.** The CH<sub>2</sub>N<sup>+</sup> ions (Table II) formed by protonation of HCN (using H<sub>2</sub>O or HCN) and by unimolecular dissociation of CH<sub>3</sub>CN and CH<sub>3</sub>NC all give the same CA mass spectrum within experimental error, indicating that they have a common structure. Of the three possible structures H<sub>2</sub>CN<sup>+</sup>, HCNH<sup>+</sup>, and CNH<sub>2</sub><sup>+</sup>, only one (HCNH<sup>+</sup>) has hydrogen atoms which are nonequivalent. CHDN<sup>+</sup> ions were made in two ways, chemical ionization of HCN using D<sub>2</sub>O as the reagent gas, and of DCN using H<sub>2</sub>O as the reagent gas. Their CA spectra (Table II) are substantially different, supporting the structure HC≡N<sup>+</sup>H, which has been predicted to be the most stable.<sup>9</sup> These CA spectra are also consistent with the expected structures HCND<sup>+</sup> and DCNH<sup>+</sup>, respectively. The m/z 13 (CH<sup>+</sup>) and 16 (DN<sup>+</sup>) are larger in the presumed HCND<sup>+</sup> spectrum, and m/z 14 (CD<sup>+</sup> + N<sup>+</sup>) and 15 (HN<sup>+</sup>) are larger in the DCNH<sup>+</sup> spectrum. Chemical ionization of CD<sub>3</sub>CN with H<sub>2</sub>O as the reagent gas gives CHDN<sup>+</sup> fragment ions whose CA spectrum resembles that of DCNH<sup>+</sup>, and chemical ionization of CH<sub>3</sub>CN with D<sub>2</sub>O gives CHDN<sup>+</sup> ions whose CA spectrum resembles the HCND<sup>+</sup> spectrum. The values of ([m/z 13<sup>+</sup>] + [16<sup>+</sup>])/([14<sup>+</sup>] + [15<sup>+</sup>]) indicate the majority of CH<sub>3</sub>C≡N<sup>+</sup>H<sup>+</sup> ions formed by the protonation of CH<sub>3</sub>CN yield HC≡N<sup>+</sup>H<sup>+</sup> ions without H-H' scrambling. Despite the mechanistic diversity of the reactions forming HCNH<sup>+</sup> from CH<sub>3</sub>CN<sup>+</sup>, CH<sub>3</sub>NC<sup>+</sup>, and CD<sub>3</sub>CNH<sup>+</sup>, and by chemical ionization, in none is there evidence for concomitant CNH<sub>2</sub><sup>+</sup> formation (such as an increased CA abundance of NH<sub>2</sub><sup>+</sup>); CNH<sub>2</sub><sup>+</sup> is predicted<sup>9</sup> to be 2.0 eV less stable than HCNH<sup>+</sup>, but to have a 1.3-eV isomerization barrier.

The CA spectra of HCND<sup>+</sup> and DCNH<sup>+</sup> also provide further evidence of the higher stability of HNC<sup>+</sup> vs. that of HCN<sup>+</sup>. For HCNH<sup>+</sup> the loss of H is ~15% greater than the loss of H' after offsetting the isotope effects.

**Losses of Neutral CH and CH<sub>2</sub>.** Peaks in electron ionization (EI) mass spectra due to the loss of CH<sub>2</sub> are unusual; for example, this accounts for a 2% peak in the mass spectrum of (C<sub>6</sub>H<sub>5</sub>)<sub>2</sub>P-CH<sub>2</sub>P(C<sub>6</sub>H<sub>5</sub>)<sub>2</sub>.<sup>7</sup> Confirming the EI spectral behavior of CH<sub>3</sub>CN and CH<sub>3</sub>NC reported above, their CA spectra show m/z 26, 27 and 28 abundances (total of m/z 12-40 = 100%) of 2.4, 2.4, and 3.8%, respectively, for CH<sub>3</sub>CN<sup>+</sup>, and 2.0, 2.4, and 3.4%, respectively, for CH<sub>3</sub>NC<sup>+</sup>. Further, the CA spectrum of CD<sub>3</sub>C≡N<sup>+</sup>H shows abundances (total of m/z 12-44 = 100%) from losses of CD<sub>2</sub>, CHD, and CD of 3.5, 2.7, and 1.8%, respectively.

Appearance energy determinations for CH<sub>3</sub>CN and CH<sub>3</sub>NC, respectively, gave CN<sup>+</sup>, 20 and 20 eV; CHN<sup>+</sup>, 15 and 15 eV; and CH<sub>2</sub>N<sup>+</sup>, 16 and 15 eV. (The photoelectron spectrum of CH<sub>3</sub>CN, ionization energy 12.2 eV, shows almost no ion formation in the range 13.2-15.1 eV.)<sup>13</sup> Based on heats of formation,<sup>14</sup> the dissociation of CH<sub>3</sub>CN<sup>+</sup> to lose CH<sub>3</sub>, CH<sub>2</sub>, and CH requires 6.9, 5.3, and 2.6 eV/mol, respectively, somewhat higher than indicated by the measured appearance energies. The favorable energies for the losses of CH<sub>2</sub> and CH vs. that of CH<sub>3</sub> more than offset the severe entropic requirements of the accompanying hydrogen rearrangements. This case is unusual because the addition of one or two hydrogens to CN<sup>+</sup> stabilizes it by 3.6 or 8.9 eV;<sup>14</sup> for comparison, the corresponding stabilizations for C<sub>2</sub>H<sub>3</sub><sup>+</sup> are 0.5 and 2.2 eV.<sup>12</sup>

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